

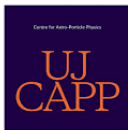
# Quantum Spread Complexity in Neutrino Oscillations

Workshop on Quantum Information Theory in Quantum Field Theory and Cosmology, BIRS, Alberta

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# Plan of this talk

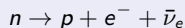
- 1 Motivation
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## Motivation

- Quantum computational complexity estimates the difficulty of constructing quantum states from elementary operations, a problem of prime importance for quantum computation.
- It can also serve to study a completely different physical problem - that of information processing inside black holes.
- Extends the connection between geometry and information.  
Growth of complexity is equal to the growth of black hole interiors.  
[\[Susskind et al., \(2014\)\]](#)
- It would be intriguing to investigate what characteristics complexity shows in other natural processes of evolution.
- Neutrinos have shown features such as entanglement and nonlocal correlations that proves their efficiency to perform QIP tasks.  
[\[blasone et al., \(2009\)\]](#), [\[Formaggio et al., \(2016\)\]](#)
- It gives us motivation to see how complex is an evolution of neutrino system and if complexity can also probe any open issue in the neutrino sector.

## Neutrino-properties

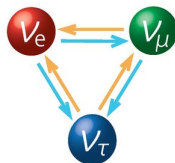
- Postulated first by Wolfgang Pauli to explain how beta decay could conserve energy, momentum and angular momentum (spin)



- Spin half, very small mass, no electric charge
- Come in three flavours  $\rightarrow \nu_e, \nu_\mu, \nu_\tau$
- Interact only via weak interaction
- Neutrinos  $\rightarrow$  Left-handed, Anti-neutrinos  $\rightarrow$  Right-handed

## Neutrino Oscillations

- About 65 billion ( $6.5 \times 10^{10}$ ) neutrinos coming from Sun's interior pass through 1 square centimeter per second. Homestake experiment's observed value was 1/3 of the predicted flux. This led Pontecorvo to suggest neutrino oscillations.
- Up-down asymmetry of atmospheric muon neutrino flux by IMB and KamioKande experiments gave additional hint of neutrino oscillations. (*T. Kajita (SK), A. McDonald (SNO), 2015*)



- Experiments : Solar (e.g. Homestake, Gallex/SAGE, SNO), Atmospheric (e.g. Super Kamiokande), Reactor (e.g. CHOOZ, KamLAND, Daya-Bay, RENO), Accelerator (e.g. T2K, MINOS,  $\text{NO}\nu\text{A}$ , DUNE (upcoming))

## Quantum mechanics in neutrino oscillations

- The three flavor states (eigenstates of weak interaction, which are detectable in lab) of neutrinos,  $\nu_e, \nu_\mu$  and  $\nu_\tau$  mix via a  $3 \times 3$  unitary matrix to form the three mass eigenstates (which are the propagation eigenstates)  $\nu_1, \nu_2$  and  $\nu_3$ . Neutrino oscillations occur only if the three corresponding masses,  $m_1, m_2$  and  $m_3$ , are non-degenerate.
- In three flavor neutrino oscillation  
 Propagation states  $\rightarrow \{|\nu_1\rangle, |\nu_2\rangle, |\nu_3\rangle\}$ ;  
 Flavor states  $\rightarrow \{|\nu_e\rangle, |\nu_\mu\rangle, |\nu_\tau\rangle\}$

- The general state of a neutrino can be expressed in flavor basis as:

$$|\Psi(t)\rangle = \nu_e(t) |\nu_e\rangle + \nu_\mu(t) |\nu_\mu\rangle + \nu_\tau(t) |\nu_\tau\rangle$$

- Same state in propagation basis looks like:

$$|\Psi(t)\rangle = \nu_1(t) |\nu_1\rangle + \nu_2(t) |\nu_2\rangle + \nu_3(t) |\nu_3\rangle$$

- The coefficients in two representations are connected by a *unitary* matrix

$$\begin{pmatrix} \nu_e(t) \\ \nu_\mu(t) \\ \nu_\tau(t) \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1(t) \\ \nu_2(t) \\ \nu_3(t) \end{pmatrix}.$$

or,

$$\nu_\alpha(t) = \mathbf{U}\nu_i(t). \quad (1)$$

## Quantum mechanics in neutrino oscillations

- A convenient parametrization for  $\mathbf{U}$  or  $U(\theta_{12}, \theta_{23}, \theta_{13}, \delta)$  is given by the PMNS matrix

$$U(\theta_{12}, \theta_{23}, \theta_{13}, \delta) = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{23}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{13}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

- where  $c_{ij} = \cos \theta_{ij}$ ,  $s_{ij} = \sin \theta_{ij}$ ,  $\theta_{ij}$  being the mixing angles and  $\delta$  the  $CP$  (Charge-Parity) violating phase.
- The mass eigenstates evolve as

$$\begin{pmatrix} \nu_1(t) \\ \nu_2(t) \\ \nu_3(t) \end{pmatrix} = \begin{pmatrix} e^{-iE_1 t} & 0 & 0 \\ 0 & e^{-iE_2 t} & 0 \\ 0 & 0 & e^{-iE_3 t} \end{pmatrix} \begin{pmatrix} \nu_1(0) \\ \nu_2(0) \\ \nu_3(0) \end{pmatrix},$$

or,

$$\nu_{\mathbf{m}}(t) = \mathbf{E} \nu_{\mathbf{m}}(0) \quad (2)$$

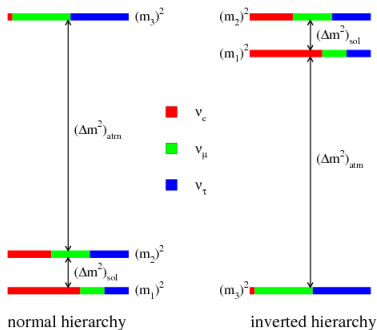
- From 1 and 2,  $\nu_{\mathbf{f}}(t) = \mathbf{U} \mathbf{E} \mathbf{U}^{-1} \nu_{\mathbf{f}}(0) = \mathbf{U}_{\mathbf{f}} \nu_{\mathbf{f}}(0)$ .

$$P_{\alpha\beta} = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2 \left( 1.27 \frac{\Delta_{ij} L}{E} \right) + 2 \sum_{i>j} \text{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin \left( 2.54 \frac{\Delta_{ij} L}{E} \right) \quad (3)$$

where  $\Delta_{ij} = m_j^2 - m_i^2 \equiv E_j - E_i$ .

## Problems not resolved yet ...

- Neutrino mass hierarchy problem *i.e.*, whether  $m_1 \leq m_2 \leq m_3$  or  $m_3 \leq m_1 \leq m_2$ ).
- CP violation ( $\delta \neq 0$ ).  
 $P(\nu_\alpha \rightarrow \nu_\beta) \neq P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta)$
- Absolute mass





## Neutrino experimental facilities

We included accelerator  $\nu_\mu$ - neutrino experimental conditions in our study such as

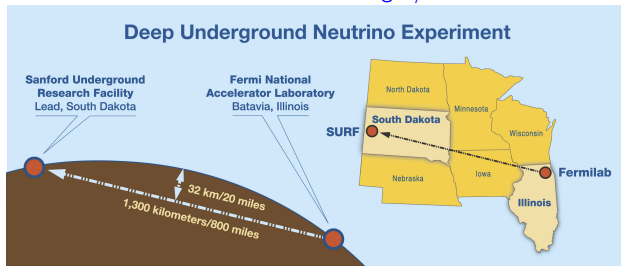
*DUNE* ( $L = 1300$  Km,  $E = 1 - 10$  GeV,  $A = 1.7 \times 10^{-13}$  eV)

*NO $\nu$ A* ( $L = 810$  Km,  $E = 1 - 4$  GeV,  $A = 1.7 \times 10^{-13}$  eV)

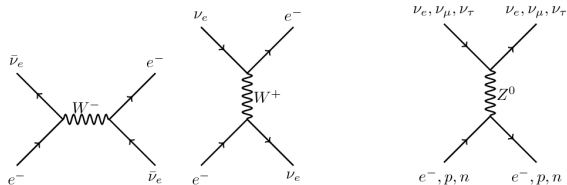
*T2K* ( $L = 295$  Km,  $E = 0.1 - 1$  GeV,  $A = 1.01 \times 10^{-13}$  eV)

( $L \rightarrow$  baseline,  $E \rightarrow$  neutrino-energy,  $A \rightarrow$  matter density potential)

Source: [www.fnal.gov/](http://www.fnal.gov/)



## Matter effects on neutrino oscillations



(a) The Feynman diagrams for charged current interactions (b) The Feynman diagram for neutral current interactions

$$H_f = UH_m U^{-1} + V \text{diag}(1, 0, 0) + V_{Z_0} \mathbb{1}_{3 \times 3}.$$

where,  $V \rightarrow$  matter density potential due to coherent-forward scattering of  $\nu_e$  with  $e^-$  present in the matter.

# Complexity

- How difficult is it to construct a desired target state with the elementary operations (gates) at your end?
- Or, the minimum number of unitaries required to construct a “target state” through a “reference state”.
- For a system  $|\phi(s)\rangle$ , if

$$U_1 U_2 U_3 U_2 |\phi(s)\rangle = U_3 U_1 U_2 U_1 (U_1)^3 U_2 |\phi(s)\rangle,$$

then the **complexity** = 4.

## Complexity of spread of states

Balasubramanian et al., PRD 106, 046007 (2022)

- The complexity of the state can be defined by minimizing the spread of the wavefunction over all possible bases.
- This minimum is uniquely attained by an orthonormal basis produced by applying the Gram-Schmidt procedure.

Schrodinger equation for a system represented by  $|\psi(t)\rangle$

$$i \frac{\partial}{\partial t} |\psi(t)\rangle = H |\psi(t)\rangle$$

Then, the time evolution of the state  $|\psi(t)\rangle$  is obtained as

$$|\psi(t)\rangle = e^{-iHt} |\psi(0)\rangle .$$

One can also write

$$|\psi(t)\rangle = \sum_{n=0}^{\infty} \frac{(-it)^n}{n!} H^n |\psi(0)\rangle = \sum_{n=0}^{\infty} \frac{(-it)^n}{n!} |\psi_n\rangle ,$$

where,  $|\psi_n\rangle = H^n |\psi(0)\rangle$ . Hence, we can see that the time evolved system-state  $|\psi(t)\rangle$  is represented as superposition of infinite  $|\psi_n\rangle$  states.

## Complexity of spread of states

We have  $|\psi_n\rangle = H^n |\psi(0)\rangle$ . These states  $\{|\psi_0\rangle, |\psi_1\rangle, |\psi_2\rangle, \dots\}$  are not orthonormalized. Gram-Schmidt procedure to obtain an ordered orthonormalized basis

$$|K_0\rangle = |\psi_0\rangle,$$

$$|K_1\rangle = |\psi_1\rangle - \frac{\langle K_0|\psi_1\rangle}{\langle K_0|K_0\rangle} |K_0\rangle,$$

$$|K_2\rangle = |\psi_2\rangle - \frac{\langle K_0|\psi_2\rangle}{\langle K_0|K_0\rangle} |K_0\rangle - \frac{\langle K_1|\psi_2\rangle}{\langle K_1|K_1\rangle} |K_1\rangle, \text{ and so on.}$$

$$\mathcal{K} = \{|K_n\rangle, n = 0, 1, 2, \dots\} \Rightarrow \text{Krylov basis}$$

**Cost function to quantify the complexity** (Balasubramanian et al., PRD 106, 046007 (2022))

For a time evolved state  $|\psi(t)\rangle$  and the Krylov basis defined as  $\{|K_n\rangle\}$ , the cost function is

$$\chi = \sum_{n=0}^{\infty} n |\langle K_n|\psi(t)\rangle|^2,$$

where  $n = 0, 1, 2, \dots$ . For such Krylov basis the above defined cost function becomes minimum.

## Spread complexity in two flavor neutrino oscillations

The evolution of flavor states can be represented by Schrodinger equation as

$$i \frac{\partial}{\partial t} \begin{pmatrix} |\nu_e(t)\rangle \\ |\nu_\mu(t)\rangle \end{pmatrix} = H_f \begin{pmatrix} |\nu_e(t)\rangle \\ |\nu_\mu(t)\rangle \end{pmatrix} \quad (4)$$

where  $H_f = UH_mU^{-1}$ ,  $U$  being the mixing matrix and  $H_m$  is the Hamiltonian (diagonal) that governs the time evolution of neutrino mass eigenstate

$$H_m = \begin{pmatrix} E_1 & 0 \\ 0 & E_2 \end{pmatrix}, \quad U = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}.$$

$$|\nu_e(0)\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad |\nu_\mu(0)\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

We have

$$\{|\psi_n\rangle\} = \begin{cases} \{|\nu_e(0)\rangle, H_f |\nu_e(0)\rangle, H_f^2 |\nu_e(0)\rangle \dots\} & \text{for initial } \nu_e \text{ flavor} \\ \{|\nu_\mu(0)\rangle, H_f |\nu_\mu(0)\rangle, H_f^2 |\nu_\mu(0)\rangle \dots\} & \text{for initial } \nu_\mu \text{ flavor} \end{cases}$$

After applying Gram-Schmidt procedure we get  $\{|\mathcal{K}_n\rangle\} = \{|\mathcal{K}_0\rangle, |\mathcal{K}_1\rangle\}$ , i.e.,

$$\{|\mathcal{K}_n\rangle\} = \begin{cases} \{|\mathcal{K}_0\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}, |\mathcal{K}_1\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}\} = \{|\nu_e\rangle, |\nu_\mu\rangle\} & \text{for initial } \nu_e \\ \{|\mathcal{K}_0\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}, |\mathcal{K}_1\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}\} = \{|\nu_\mu\rangle, |\nu_e\rangle\} & \text{for initial } \nu_\mu \end{cases}$$

## Spread complexity in two flavor neutrino oscillations

For a time evolved state  $|\nu_e(t)\rangle = \begin{pmatrix} A_{ee}(t) \\ A_{e\mu}(t) \end{pmatrix} = \begin{pmatrix} \cos^2 \theta e^{-iE_1 t} + \sin^2 \theta e^{-iE_2 t} \\ \sin \theta \cos \theta (e^{-iE_2 t} - e^{-iE_1 t}) \end{pmatrix}$   
 (with  $\{|K_n\rangle\} = \{|\nu_e(0)\rangle, |\nu_\mu(0)\rangle\}$ )

$$\chi_e = \sum_{n=0}^1 n |\langle K_n | \nu_e(t) \rangle|^2 = P_{e\mu}$$

Similarly, for state  $|\nu_\mu(t)\rangle = (A_{\mu e}(t), A_{\mu\mu}(t))^T$  (with  $\{|K_n\rangle\} = \{|\nu_\mu(0)\rangle, |\nu_e(0)\rangle\}$ )

$$\chi_\mu = P_{\mu e}$$

- The more the oscillation probability of neutrino flavor, the more complex the evolution of the neutrino flavor state.
- Since  $P_{e\mu} = P_{\mu e}$  for standard vacuum oscillations, the complexity embedded in this system comes out to be same for both cases of initial flavor, *i.e.*, complexity of the system doesn't depend on the initial flavor of neutrino.

## Spread complexity in three flavor neutrino oscillations

We have three types of initial states as  $|\nu_e\rangle = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$ ,  $|\nu_\mu\rangle = \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix}$ ,  $|\nu_\tau\rangle = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$  with Hamiltonian  $H_f = UH_mU^{-1}$ ,  $H_m = \text{diag}(0, \Delta m_{21}^2, \Delta m_{31}^2)$  and  $U \rightarrow 3 \times 3$  PMNS mixing matrix

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{13}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

Here, Krylov basis  $\neq$  flavor basis.

- For initial  $|\nu_e\rangle$  state  $|K_0\rangle \equiv |\nu_e\rangle = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$ , other states spanning the Krylov basis take the form

$$|K_1\rangle = N_1 \begin{pmatrix} 0 \\ a_1 \\ a_2 \end{pmatrix} = N_1 \begin{pmatrix} 0 \\ \left(\frac{\Delta m_{21}^2}{2E}\right) U_{e2}^* U_{\mu 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) U_{e3}^* U_{\mu 3} \\ \left(\frac{\Delta m_{21}^2}{2E}\right) U_{e2}^* U_{\tau 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) U_{e3}^* U_{\tau 3} \end{pmatrix},$$

$$|K_2\rangle = N_2 \begin{pmatrix} 0 \\ b_1 \\ b_2 \end{pmatrix} = N_2 \begin{pmatrix} 0 \\ \left(\frac{\Delta m_{21}^2}{2E}\right) \left(\frac{\Delta m_{21}^2}{2E} - A\right) U_{e2}^* U_{\mu 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) \left(\frac{\Delta m_{31}^2}{2E} - A\right) U_{e3}^* U_{\mu 3} \\ \left(\frac{\Delta m_{21}^2}{2E}\right) \left(\frac{\Delta m_{21}^2}{2E} - A\right) U_{e2}^* U_{\tau 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) \left(\frac{\Delta m_{31}^2}{2E} - A\right) U_{e3}^* U_{\tau 3} \end{pmatrix}$$



## Spread complexity in three flavor neutrino oscillations

$$\chi_e = P_{e\mu}(t)(N_1^2|a_1|^2 + 2N_2^2|b_1|^2) + P_{e\tau}(t)(N_1^2|a_2|^2 + 2N_2^2|b_2|^2) + 2\Re(N_1^2 a_1^* a_2 A_{e\mu}(t) A_{e\tau}(t)^*) + 4\Re(N_2^2 b_1^* b_2 A_{e\mu}(t) A_{e\tau}(t)^*)$$

with

$$A = \frac{\left( (\Delta m_{21}^2)^3 |U_{\alpha 2}|^2 (1 - |U_{\alpha 2}|^2) + (\Delta m_{31}^2)^3 |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) - (\Delta m_{21}^2) (\Delta m_{31}^2) |U_{\alpha 2}|^2 |U_{\alpha 3}|^2 (\Delta m_{21}^2 + \Delta m_{31}^2) \right)}{(\Delta m_{21}^2)^2 |U_{\alpha 2}|^2 (1 - |U_{\alpha 2}|^2) + (\Delta m_{31}^2)^2 |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) - 2 (\Delta m_{21}^2) (\Delta m_{31}^2) |U_{\alpha 2}|^2 |U_{\alpha 3}|^2},$$

## Effects of different oscillation parameters

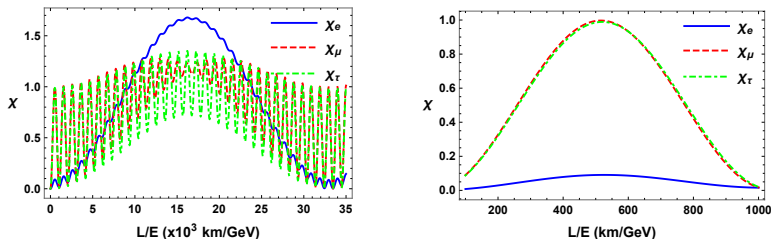
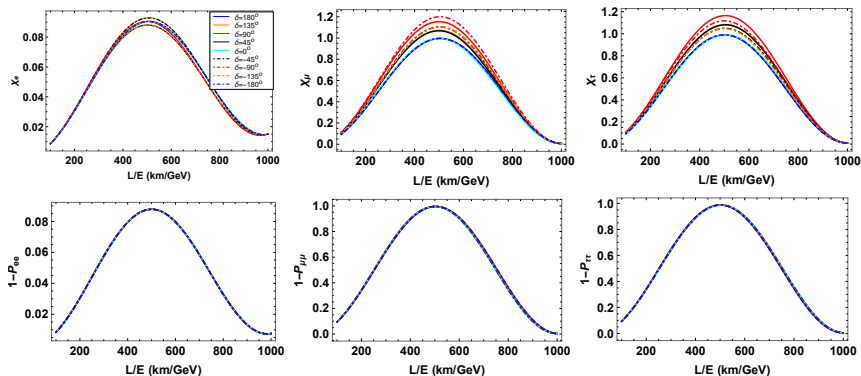


Figure: Complexity plotted with respect to the distance  $L$  over energy  $E$  traveled by neutrinos in vacuum and in case if the initial flavor is  $\nu_e$  (blue solid line),  $\nu_\mu$  (red dashed line) and  $\nu_\tau$  (green dot-dashed line) for  $CP$ -violating phase  $\delta = 0^\circ$ . Here, mixing parameters  $\theta_{12} = 33.64^\circ$ ,  $\theta_{13} = 8.53^\circ$ ,  $\theta_{23} = 47.63^\circ$ ,  $\Delta m_{21}^2 = 7.53 \times 10^{-5} \text{ eV}^2$  and  $\Delta m_{31}^2 = 2.45 \times 10^{-3} \text{ eV}^2$  are considered.

- The rapid oscillation pattern seen in the left panel (zoomed-in in the right panel) is due to  $\Delta m_{31}^2$  mass-squared difference in the oscillation phase, while the longer oscillation pattern is due to  $\Delta m_{21}^2$  in the oscillation phase. The oscillation length is  $\sim 10^3 \text{ km}$  at  $E = 1 \text{ GeV}$  for  $\Delta m_{31}^2$  and  $\sim 3 \times 10^4 \text{ km}$  at  $E = 1 \text{ GeV}$  for  $\Delta m_{21}^2$ .
- In the general case the complexity is maximum if the neutrino is produced initially as  $\nu_e$ , however, this happens only at a very large  $L/E$  value of  $\sim 1.6 \times 10^4 \text{ km/GeV}$ .
- In current experimental setups (right panel), which covers roughly one oscillation length for  $\Delta m_{31}^2$ , the initial  $\nu_e$  flavor provides the least complexity among all neutrino flavors.

Effects of  $CP$ -violating parameter  $\delta$ Figure: Complexities and  $1 - P_{\alpha\alpha}$  with respect to  $L/E$ .

- Complexity mimics the features of the total oscillation probability  $1 - P_{\alpha\alpha}$ .
- However, it is visible that  $\chi_\alpha$  for all three flavors provide more information regarding the  $CP$ -violating phase  $\delta$ .

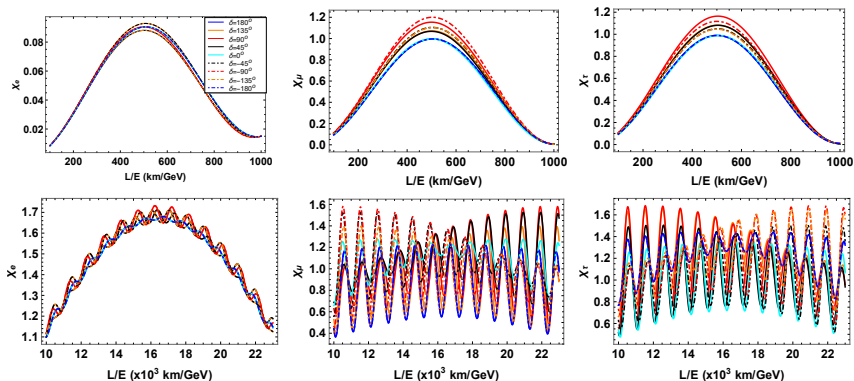


Figure: Complexity for small  $L/E$  range (upper panels), large  $L/E$  range (lower panels) with respect to  $L/E$  for initial flavor is  $\nu_e$  (left),  $\nu_\mu$  (middle) and  $\nu_\tau$  (right) for different values of the  $CP$ -phase  $\delta$  depicted by different colors.

- For large  $L/E$  range the complexities are maximized and the corresponding  $\delta = +90^\circ$  or  $-90^\circ$  for  $\chi_\mu$  and  $\chi_\tau$ , and at  $\delta = \pm 90^\circ$  for  $\chi_e$  where  $CP$  is maximally violated.
- In the limited  $L/E$  range  $\chi_\mu$  and  $\chi_\tau$  are maximized at  $\delta = -90^\circ$  (red-dashed line) and at  $\delta = +90^\circ$  (red-solid line), respectively. However,  $\chi_e$  is maximized at  $\delta = +135^\circ$  and at  $-45^\circ$ .

## Matter effects on complexity

- For any initial flavor  $\nu_\alpha$

$$\begin{aligned} |K_0\rangle_\alpha^{matter} &= |K_0\rangle_\alpha^{vacuum} \\ |K_1\rangle_\alpha^{matter} &= |K_1\rangle_\alpha^{vacuum}, \end{aligned}$$

- $|K_2\rangle$  contains the effects of constant matter density

$$|K_2\rangle_e = N_{2e}^m(0, b_1^m, b_2^m)^T$$

where,

$$b_1^m = \left(\frac{\Delta m_{21}^2}{2E}\right) \left(\frac{\Delta m_{21}^2}{2E} + V - B_e\right) U_{e2}^* U_{\mu 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) \left(\frac{\Delta m_{31}^2}{2E} + V - B_e\right) U_{e3}^* U_{\mu 3},$$

$$b_2^m = \left(\frac{\Delta m_{21}^2}{2E}\right) \left(\frac{\Delta m_{21}^2}{2E} + V - B_e\right) U_{e2}^* U_{\tau 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) \left(\frac{\Delta m_{31}^2}{2E} + V - B_e\right) U_{e3}^* U_{\tau 3}.$$

## Matter effects on complexity

- Similarly, for the initial  $\nu_\mu$  flavor

$$|K_2\rangle_\mu = N_{2\mu}^m (d_1^m, 0, d_2^m)^T,$$

where,

$$d_1^m = \left( \frac{\Delta m_{21}^2}{2E} \right) \left( \frac{\Delta m_{21}^2}{2E} + V - B_\mu \right) U_{e2} U_{\mu 2}^* + \left( \frac{\Delta m_{31}^2}{2E} \right) \left( \frac{\Delta m_{31}^2}{2E} + V - B_\mu \right) U_{e3} U_{\mu 3}^*$$

$$d_2^m = \left( \frac{\Delta m_{21}^2}{2E} \right) \left( \frac{\Delta m_{21}^2}{2E} - B_\mu \right) U_{\mu 2}^* U_{\tau 2} + \left( \frac{\Delta m_{31}^2}{2E} \right) \left( \frac{\Delta m_{31}^2}{2E} - B_\mu \right) U_{\mu 3}^* U_{\tau 3},$$

## Matter effects on complexity

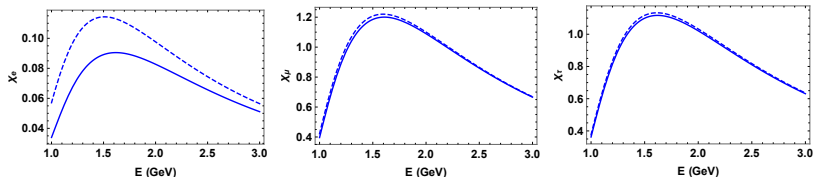


Figure: Cost function  $\chi_e$  (left),  $\chi_\mu$  (middle) and  $\chi_\tau$  (right) w. r. t. neutrino-energy  $E$  is shown. Here,  $L = 810$  km,  $\delta = -90^\circ$  and higher octant of  $\theta_{23}$  is considered. Solid and dashed curves represent the case of vacuum and matter oscillations, respectively.  $V = 1.01 \times 10^{-13}$  eV.

- Matter effect increases complexity of the system in all cases of initial flavors of the neutrino, most significantly for  $\nu_e$  as expected.

## Spread complexity in neutrino oscillation experiments

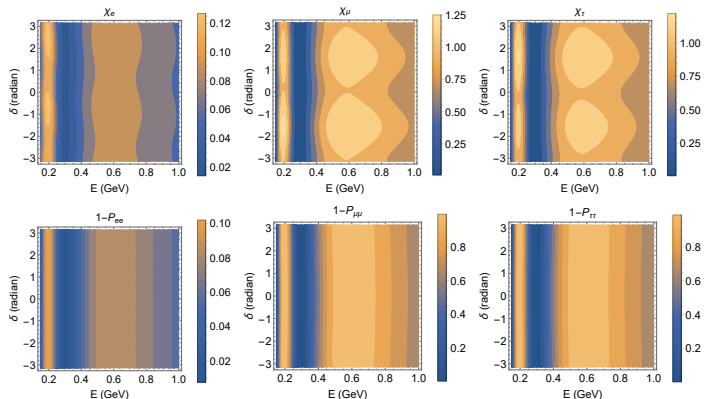


Figure: T2K: Cost function (upper panel) and  $1 - P_{\alpha\alpha}$  (lower panel) in the plane of  $E - \delta$  in case of initial flavor  $\nu_e$  (left),  $\nu_\mu$  (middle) and  $\nu_\tau$  (right). Here,  $L = 295$  km and mixing parameters  $\theta_{12} = 33.64^\circ$ ,  $\theta_{13} = 8.53^\circ$ ,  $\theta_{23} = 47.63^\circ$ ,  $\Delta m_{21}^2 = 7.53 \times 10^{-5} \text{ eV}^2$  and  $\Delta m_{31}^2 = 2.45 \times 10^{-3} \text{ eV}^2$  are considered.



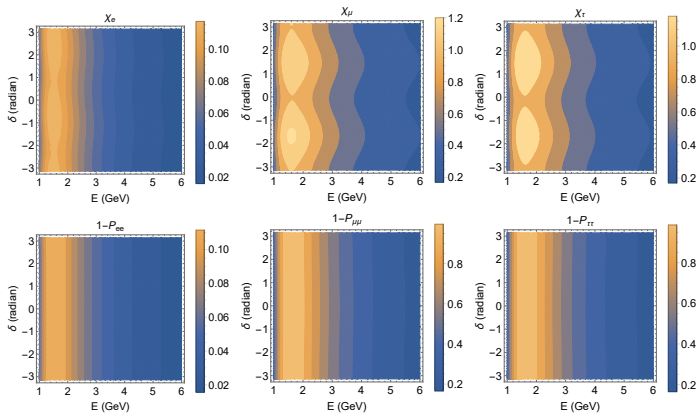


Figure: NO $\nu$ A: Cost function (upper panel) and  $1 - P_{\alpha\alpha}$  (lower panel) in the plane of  $E - \delta$  in case of initial flavor  $\nu_e$  (left),  $\nu_\mu$  (middle) and  $\nu_\tau$  (right). Here,  $L = 810$  km, and higher octant of  $\theta_{23}$  ( $47.63^\circ$ ) is considered.

- For both the experiments, the maxima of  $\chi_\mu$  and  $\chi_\tau$  are found at  $\delta \approx -\pi/2$  and  $\delta = \pi/2$ , respectively.
- This means that the matter effect just enhances the magnitude of complexities, however, the characteristics of  $\chi_\alpha$  with respect to  $\delta$  are almost similar for both T2K and NO $\nu$ A experiments.

- In the T2K and NOvA experimental setups, where only  $\nu_\mu$  beams are produced, the only relevant complexity is  $\chi_\mu$ .
- For both the T2K and NOvA  $\chi_\mu$  is maximized at  $\delta \approx -1.5$  radian at the relevant experimental energies. The T2K best-fit value of  $\delta = -2.14^{+0.90}_{-0.69}$  radian is consistent with this expectation.
- The NOvA best-fit, however, is at  $\delta \approx 2.58$  radian which is far away from the maximum  $\chi_\mu$  in the lower-half plane of  $\delta$  but is still within a region of high  $\chi_\mu$  value in the upper-half plane of  $\delta$ .
- $P_{\mu e}$ , which is the only oscillation probability accessible to the T2K and NOvA setups, it becomes maximum at  $\delta \approx -1.5$  radian. This is compatible with T2K best-fit but is in odd with the NOvA best-fit.
- Complexity provides correct prediction for the  $\delta$  in experimental setups.

## Effects of neutrino mass ordering

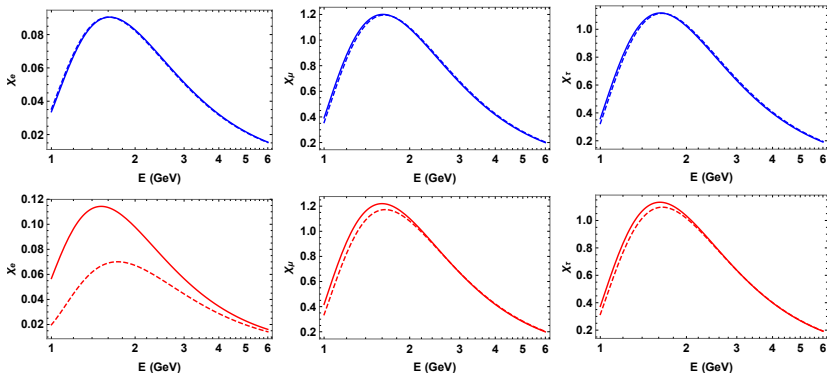


Figure: NOvA: Complexity with respect to neutrino-energy  $E$  in case of initial flavor  $\nu_e$  (left),  $\nu_\mu$  (middle) and  $\nu_\tau$  (right) with  $L = 810$  km and  $\delta = -90^\circ$ . The upper and lower panel represent the case of vacuum and matter oscillations, respectively. Solid curves are associated with NH and dashed curves depict the IH.

- Complexity can distinguish between the effects due to normal (NH) ( $+\Delta_{31}$ ) and inverted hierarchy (IH) ( $-\Delta_{31}$ ) in the presence of non-zero matter potential.

## Summary & Conclusions

- We examined the spread complexity of neutrino states in two- and three-flavor oscillation scenarios.
- In the two-flavor scenario, complexity and transition probabilities yield equivalent information.
- In case of three-flavor oscillation, initial flavor state evolves into two mixed final states. Hence, the complexity contains additional information regarding open issues related to neutrinos, compared to the total oscillation probability.
- Remarkably, we found that the complexity is maximized for a value of the phase angle for which CP is also maximally violated. T2K experimental data also favors this phase angle, which is obtained from flavor transition.
- Quantum spread complexity emerges as a potent and novel quantity for investigating neutrino oscillations. It successfully reproduces existing results, also demonstrates the potential to serve as a theoretical tool for predicting new outcomes in future experiments.

Thank you for your attention!

## BACKUP SLIDES

## Comparing the effects of neutrino mass ordering for neutrinos & antineutrinos

For antineutrino  $\rightarrow \{V \rightarrow -V, \delta \rightarrow -\delta\}$

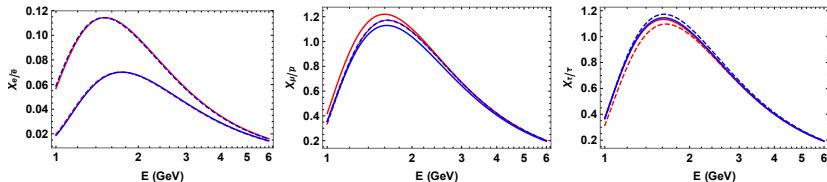


Figure: NOvA: Complexities and  $P_{\mu e}$  with respect to neutrino-energy  $E$  where red and blue curves represent neutrino and antineutrino case, respectively, with solid (normal ordering) and dashed (inverted ordering) lines. Here  $L = 810$  km and  $\delta = -90^\circ$  are considered.

- For both neutrino and antineutrino, the effects of NH and IH are significantly distinguishable for all three flavors.
- In case of  $\chi_e$ , red-solid line (neutrinos for NH) and blue-dashed line (antineutrinos for IH) exhibit more complexity, *i.e.*, complete swap between the NH (IH) hierarchy and  $\nu$  ( $\bar{\nu}$ ).
- For  $\chi_\mu$  and  $\chi_\tau$  the maximum is achieved in case of neutrinos with NH and Antineutrinos with IH, respectively.

## Spread complexity in three flavor (vacuum) neutrino oscillations

- Similarly, for initial  $|\nu_\mu\rangle$ ,  $|K_0\rangle \equiv |\nu_\mu\rangle = (0, 1, 0)^T$ , then we get

$$|K_1\rangle = N_{1\mu} \begin{pmatrix} c_1 \\ 0 \\ c_2 \end{pmatrix} = N_{1\mu} \begin{pmatrix} \left(\frac{\Delta m_{21}^2}{2E}\right) U_{\mu 2}^* U_{e2} + \left(\frac{\Delta m_{31}^2}{2E}\right) U_{\mu 3}^* U_{e3} \\ 0 \\ \left(\frac{\Delta m_{21}^2}{2E}\right) U_{\mu 2}^* U_{\tau 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) U_{\mu 3}^* U_{\tau 3} \end{pmatrix},$$

$$|K_2\rangle = N_{2\mu} \begin{pmatrix} d_1 \\ 0 \\ d_2 \end{pmatrix} = N_{2\mu} \begin{pmatrix} \left(\frac{\Delta m_{21}^2}{2E}\right) \left(\frac{\Delta m_{21}^2}{2E} - A\right) U_{\mu 2}^* U_{e2} + \left(\frac{\Delta m_{31}^2}{2E}\right) \left(\frac{\Delta m_{31}^2}{2E} - A\right) U_{\mu 3}^* U_{e3} \\ 0 \\ \left(\frac{\Delta m_{21}^2}{2E}\right) \left(\frac{\Delta m_{21}^2}{2E} - A\right) U_{\mu 2}^* U_{\tau 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) \left(\frac{\Delta m_{31}^2}{2E} - A\right) U_{\mu 3}^* U_{\tau 3} \end{pmatrix}$$

$$\begin{aligned} \chi_\mu = & P_{\mu e}(t)(N_{1\mu}^2 |c_1|^2 + 2N_{2\mu}^2 |d_1|^2) + P_{\mu \tau}(t)(N_{1\mu}^2 |c_2|^2 + 2N_{2\mu}^2 |d_2|^2) + 2\Re(N_{1\mu}^2 c_1^* c_2 A_{\mu e}(t) A_{\mu \tau}(t)^*) \\ & + 4\Re(N_{2\mu}^2 d_1^* d_2 A_{\mu e}(t) A_{\mu \tau}(t)^*). \end{aligned}$$



## Spread complexity in three flavor neutrino oscillations

- In case of  $|K_0\rangle \equiv |\nu_\tau\rangle = (0, 0, 1)^T$ ,

$$|K_1\rangle = N_{1\tau}(e_1, e_2, 0)^T = N_{1\tau} \begin{pmatrix} \left(\frac{\Delta m_{21}^2}{2E}\right) U_{\tau 2}^* U_{e2} + \left(\frac{\Delta m_{31}^2}{2E}\right) U_{\tau 3}^* U_{e3} \\ \left(\frac{\Delta m_{21}^2}{2E}\right) U_{\tau 2}^* U_{\mu 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) U_{\tau 3}^* U_{\mu 3} \\ 0 \end{pmatrix},$$

$$|K_2\rangle = N_{2\tau}(f_1, f_2, 0)^T = N_{2\tau} \begin{pmatrix} \left(\frac{\Delta m_{21}^2}{2E}\right) \left(\frac{\Delta m_{21}^2}{2E} - A\right) U_{\tau 2}^* U_{e2} + \left(\frac{\Delta m_{31}^2}{2E}\right) \left(\frac{\Delta m_{31}^2}{2E} - A\right) U_{\tau 3}^* U_{e3} \\ \left(\frac{\Delta m_{21}^2}{2E}\right) \left(\frac{\Delta m_{21}^2}{2E} - A\right) U_{\tau 2}^* U_{\mu 2} + \left(\frac{\Delta m_{31}^2}{2E}\right) \left(\frac{\Delta m_{31}^2}{2E} - A\right) U_{\tau 3}^* U_{\mu 3} \\ 0 \end{pmatrix}.$$

$$\begin{aligned} \chi_\tau = P_{\tau e}(t)(N_1^2|e_1|^2 + 2N_2^2|f_1|^2) + P_{\tau\mu}(t)(N_1^2|e_2|^2 + 2N_2^2|f_2|^2) + 2\Re(N_1^2 e_1^* e_2 A_{\tau e}(t) A_{\tau\mu}(t)^*) \\ + 4\Re(N_2^2 f_1^* f_2 A_{\tau e}(t) A_{\tau\mu}(t)^*). \end{aligned}$$

Here,

$$A = \frac{\left[ \begin{aligned} &(\Delta m_{21}^2)^3 |U_{\alpha 2}|^2(1 - |U_{\alpha 2}|^2) + (\Delta m_{31}^2)^3 |U_{\alpha 3}|^2(1 - |U_{\alpha 3}|^2) \\ &- (\Delta m_{21}^2) (\Delta m_{31}^2) |U_{\alpha 2}|^2 |U_{\alpha 3}|^2 (\Delta m_{21}^2 + \Delta m_{31}^2) \end{aligned} \right]}{(\Delta m_{21}^2)^2 |U_{\alpha 2}|^2(1 - |U_{\alpha 2}|^2) + (\Delta m_{31}^2)^2 |U_{\alpha 3}|^2(1 - |U_{\alpha 3}|^2) - 2 (\Delta m_{21}^2) (\Delta m_{31}^2) |U_{\alpha 2}|^2 |U_{\alpha 3}|^2},$$

## Spread complexity in three flavor neutrino oscillations

$$N_{1\alpha} = \left( \left( \frac{\Delta m_{21}^2}{2E} \right)^2 |U_{\alpha 2}|^2 (1 - |U_{\alpha 2}|^2) + \left( \frac{\Delta m_{31}^2}{2E} \right)^2 |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) - 2 \left( \frac{\Delta m_{21}^2}{2E} \right) \left( \frac{\Delta m_{31}^2}{2E} \right) |U_{\alpha 2}|^2 |U_{\alpha 3}|^2 \right)^{-1/2},$$

$$N_{2\alpha} = \left( \left( \frac{\Delta m_{21}^2}{2E} \right)^2 \left( \frac{\Delta m_{21}^2}{2E} - A \right)^2 |U_{\alpha 2}|^2 (1 - |U_{\alpha 2}|^2) + \left( \frac{\Delta m_{31}^2}{2E} \right)^2 \left( \frac{\Delta m_{31}^2}{2E} - A \right)^2 |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) - 2 \left( \frac{\Delta m_{21}^2}{2E} \right) \left( \frac{\Delta m_{31}^2}{2E} \right) \left( \frac{\Delta m_{21}^2}{2E} - A \right) \left( \frac{\Delta m_{31}^2}{2E} - A \right) |U_{\alpha 2}|^2 |U_{\alpha 3}|^2 \right)^{-1/2}$$

## Matter effects on complexity

$$B_e = \left[ \left( \Delta m_{21}^2 \right)^2 \left( \Delta m_{21}^2 + 2EV \right) |U_{e2}|^2 (1 - |U_{e2}|^2) + \left( \Delta m_{31}^2 \right)^2 \left( \Delta m_{31}^2 + 2EV \right) |U_{e3}|^2 \right. \\ \left. (1 - |U_{e3}|^2) - \left( \Delta m_{21}^2 \right) \left( \Delta m_{31}^2 \right) |U_{e2}|^2 |U_{e3}|^2 \left( \left( \Delta m_{21}^2 + 2EV \right) + \left( \Delta m_{31}^2 + 2EV \right) \right) \right] \\ \left[ 2E \left[ \left( \Delta m_{21}^2 \right)^2 |U_{e2}|^2 (1 - |U_{e2}|^2) + \left( \Delta m_{31}^2 \right)^2 |U_{e3}|^2 (1 - |U_{e3}|^2) \right. \right. \\ \left. \left. - 2 \left( \Delta m_{21}^2 \right) \left( \Delta m_{31}^2 \right) |U_{e2}|^2 |U_{e3}|^2 \right] \right]^{-1}.$$

For initial  $\nu_\mu$  and  $\nu_\tau$  state the constant  $B_\alpha$  is

$$B_\alpha = \left[ \left( \Delta m_{21}^2 \right)^3 |U_{\alpha 2}|^2 (1 - |U_{\alpha 2}|^2) + \left( \Delta m_{31}^2 \right)^3 |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) - \left( \Delta m_{21}^2 \right) \left( \Delta m_{31}^2 \right) \right. \\ \left. |U_{\alpha 2}|^2 |U_{\alpha 3}|^2 \left( \Delta m_{21}^2 + \Delta m_{31}^2 \right) + 2EV \left( \left( \Delta m_{21}^2 \right)^2 |U_{e2}|^2 |U_{\alpha 2}|^2 + \left( \Delta m_{31}^2 \right)^2 |U_{e3}|^2 |U_{\alpha 3}|^2 \right. \right. \\ \left. \left. + 2 \left( \Delta m_{21}^2 \right) \left( \Delta m_{31}^2 \right) \Re(U_{e2}^* U_{\alpha 2} U_{e3} U_{\alpha 3}^*) \right) \right] \left[ 2E \left[ \left( \Delta m_{21}^2 \right)^2 |U_{\alpha 2}|^2 (1 - |U_{\alpha 2}|^2) + \left( \Delta m_{31}^2 \right)^2 \right. \right. \\ \left. \left. |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) - 2 \left( \Delta m_{21}^2 \right) \left( \Delta m_{31}^2 \right) |U_{\alpha 2}|^2 |U_{\alpha 3}|^2 \right] \right]^{-1},$$

## Matter effects on complexity

$$\begin{aligned}
 N_{2e}^m = & \left( \left( \frac{\Delta m_{21}^2}{2E} \right)^2 |U_{e2}|^2 (1 - |U_{e2}|^2) \left[ \left( \frac{\Delta m_{21}^2}{2E} + V - B_e \right)^2 \right] \right. \\
 & + \left. \left( \frac{\Delta m_{31}^2}{2E} \right)^2 |U_{e3}|^2 (1 - |U_{e3}|^2) \left[ \left( \frac{\Delta m_{31}^2}{2E} + V - B_e \right)^2 \right] \right. \\
 & \left. - 2 \left( \frac{\Delta m_{21}^2}{2E} \right) \left( \frac{\Delta m_{31}^2}{2E} \right) \left( \frac{\Delta m_{21}^2}{2E} + V - B_e \right) \left( \frac{\Delta m_{31}^2}{2E} + V - B_e \right) |U_{e2}|^2 |U_{e3}|^2 \right)^{-1/2},
 \end{aligned}$$

$$\begin{aligned}
 N_{2\mu}^m = & \left( \left( \frac{\Delta m_{21}^2}{2E} \right)^2 |U_{\mu 2}|^2 \left[ \left( \frac{\Delta m_{21}^2}{2E} + V - B_\mu \right)^2 |U_{e2}|^2 + \left( \frac{\Delta m_{21}^2}{2E} - B_\mu \right)^2 |U_{\tau 2}|^2 \right] \right. \\
 & + \left. \left( \frac{\Delta m_{31}^2}{2E} \right)^2 |U_{\mu 3}|^2 \left[ \left( \frac{\Delta m_{31}^2}{2E} + V - B_\mu \right)^2 |U_{e3}|^2 + \left( \frac{\Delta m_{31}^2}{2E} - B_\mu \right)^2 |U_{\tau 3}|^2 \right] \right. \\
 & + 2 \left( \frac{\Delta m_{21}^2}{2E} \right) \left( \frac{\Delta m_{31}^2}{2E} \right) \left[ \left( \frac{\Delta m_{21}^2}{2E} + V - B_\mu \right) \left( \frac{\Delta m_{31}^2}{2E} + V - B_\mu \right) \Re(U_{\mu 2}^* U_{e2} U_{\mu 3} U_{e3}^*) \right. \\
 & \left. + \left( \frac{\Delta m_{21}^2}{2E} - B_\mu \right) \left( \frac{\Delta m_{31}^2}{2E} - B_\mu \right) \Re(U_{\mu 2}^* U_{\tau 2} U_{\mu 3} U_{\tau 3}^*) \right] \right)^{-1/2},
 \end{aligned}$$

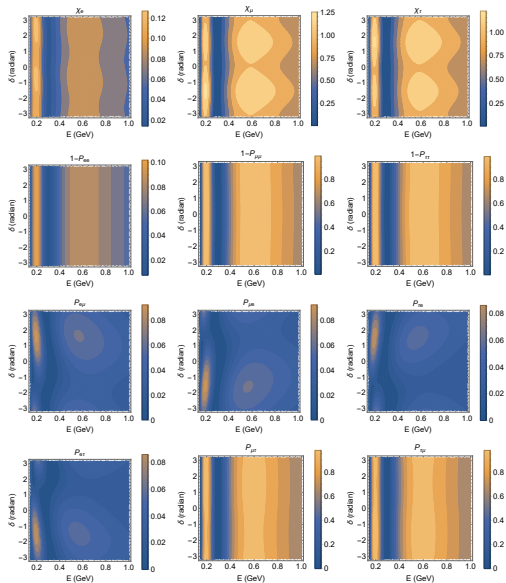


Figure: T2K: Cost function (upper panel) and  $1 - P_{\alpha\alpha}$  (lower panel) in the plane of  $E - \delta$  in case of initial flavor  $\nu_e$  (left),  $\nu_\mu$  (middle) and  $\nu_\tau$  (right). Here,  $L = 295$  km is considered.